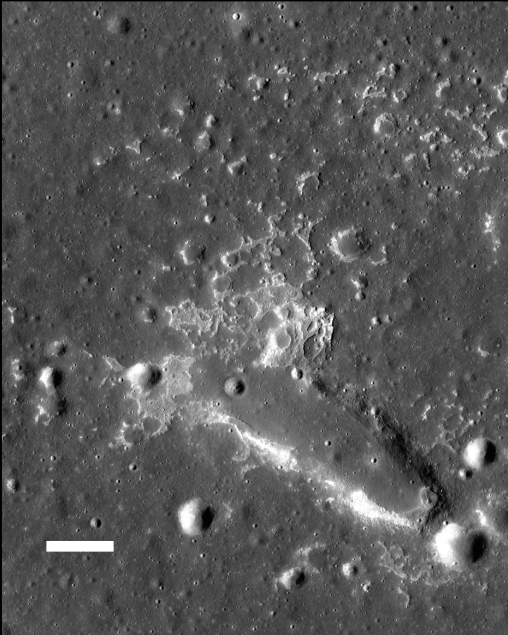


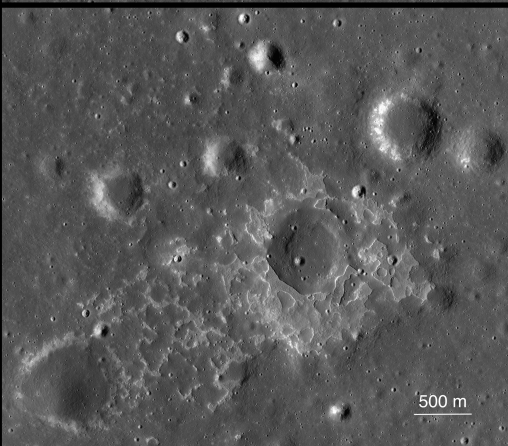
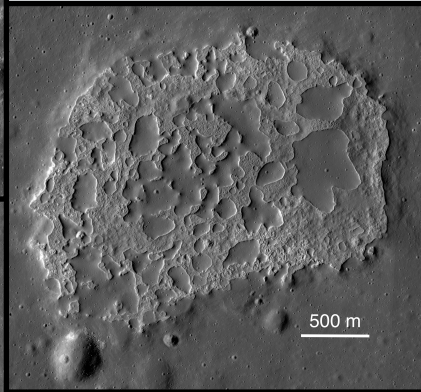
Low Thermal Inertia Volcanic Deposits on the Moon

Catherine M. Elder¹, Paul O. Hayne¹, and Tim D.
Glotch²

¹Jet Propulsion Laboratory, California Institute of Technology; ²Stony
Brook University



Irregular
mare
patches



Pyroclastic deposits

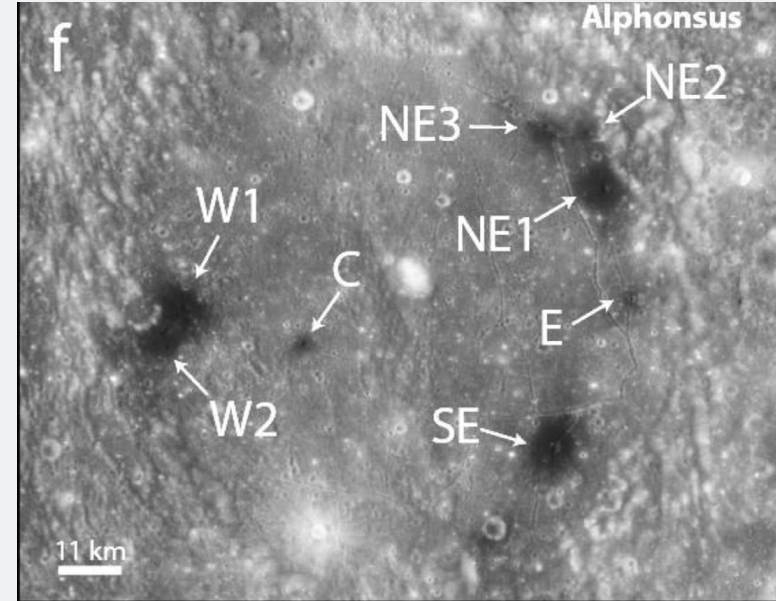
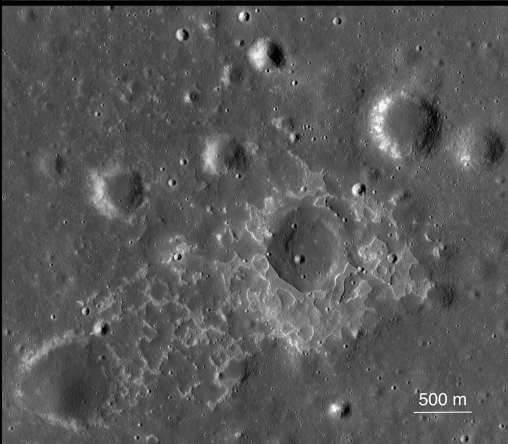
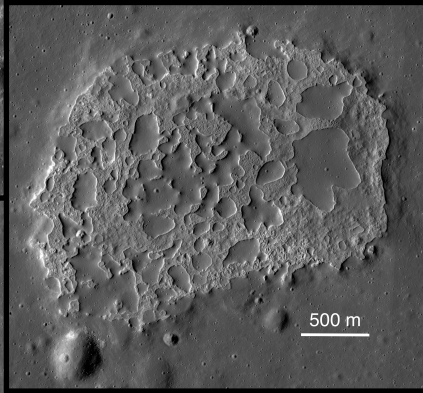


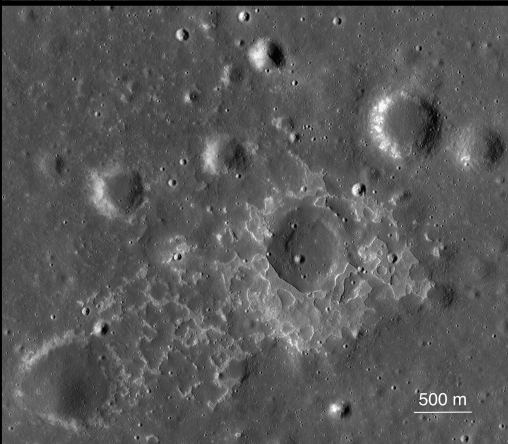
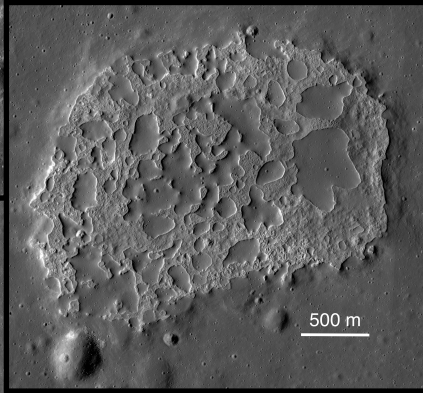
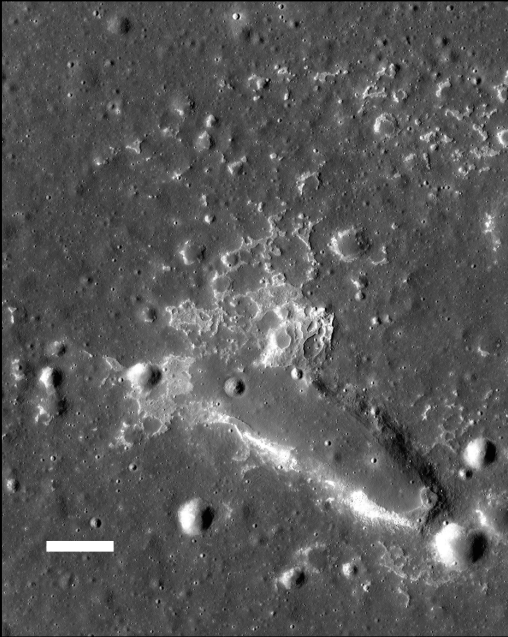
Figure credit: Trang et al. (2017)

Irregular mare patches



- Small 100 – 5000 m
 - 70 identified
 - Young? (controversial)
 - Crater counting
 - Optical maturity
 - Topographic relief
- Inconsistent with thermal evolution models and other examples of lunar volcanism

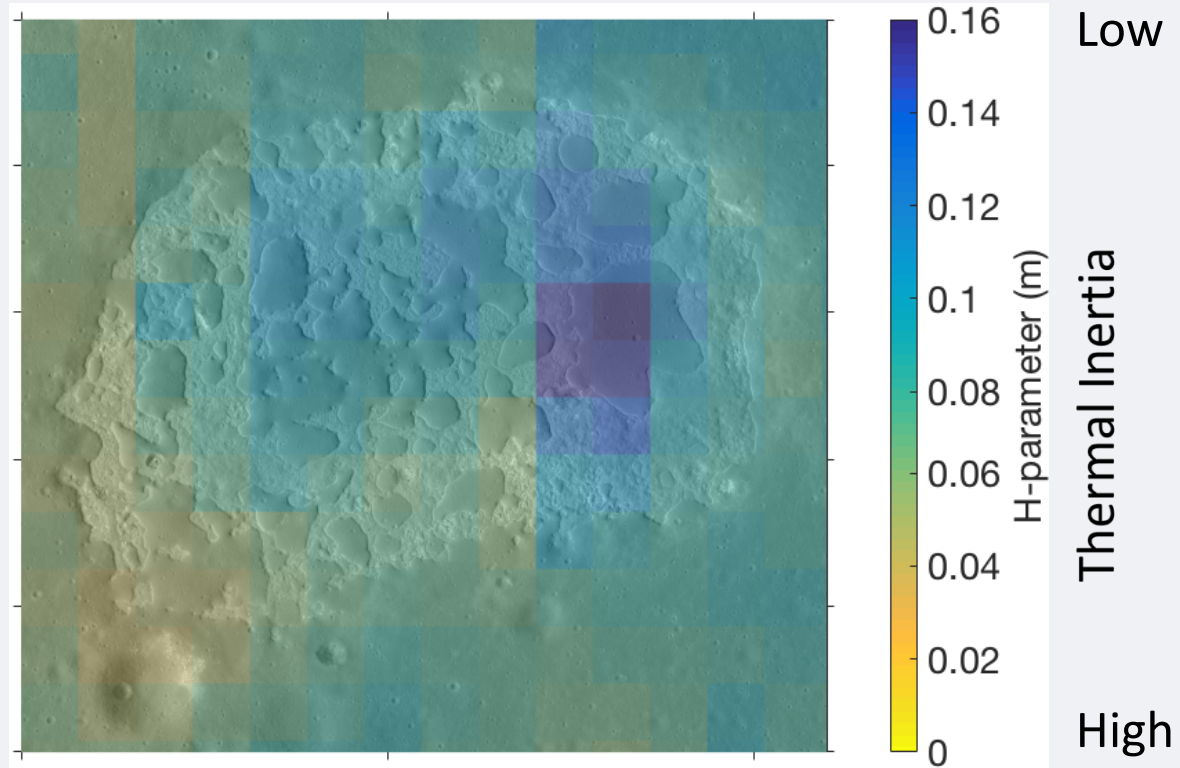
Irregular mare patches



- Caldera Collapse (El-Baz, 1973)
- Explosive outgassing (Schultz et al., 2006)
- Lava flow inflation (Garry et al., 2012)
- Pyroclastic eruptions (Carter et al., 2013)
- Regolith drainage (Qiao et al., 2016)
- Vesicular magmatic foam (Wilson et al., 2017)

Irregular Mare Patches

- Data from the Lunar Reconnaissance Orbiter (LRO) Diviner radiometer
- Ina has a low thermal inertia (Elder et al., 2017)
- Pyroclastic eruptions? Magmatic foam?



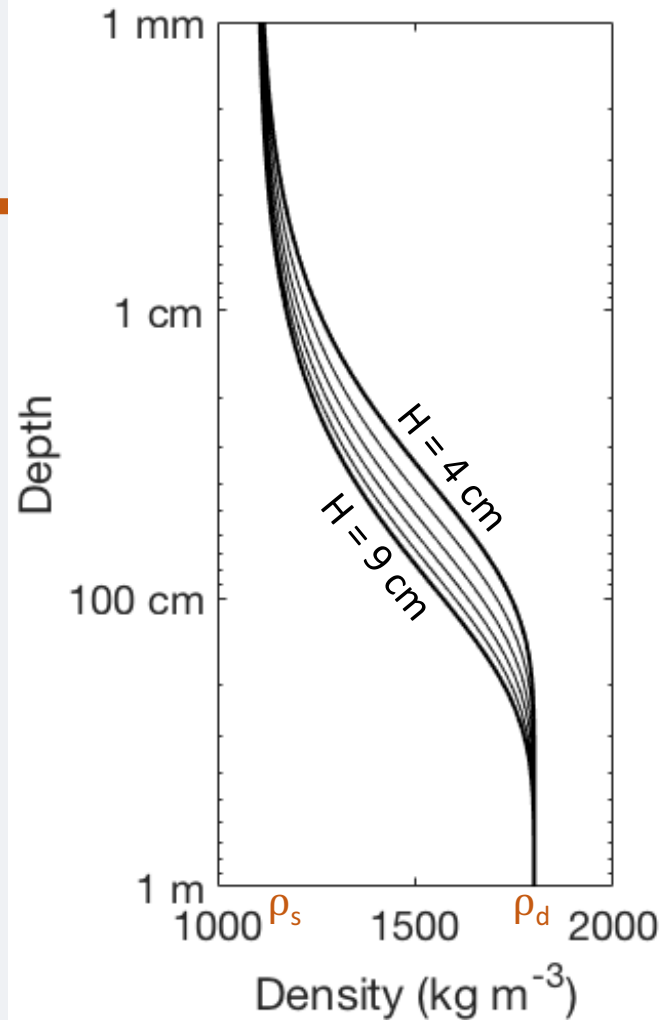
Pyroclastic Deposits

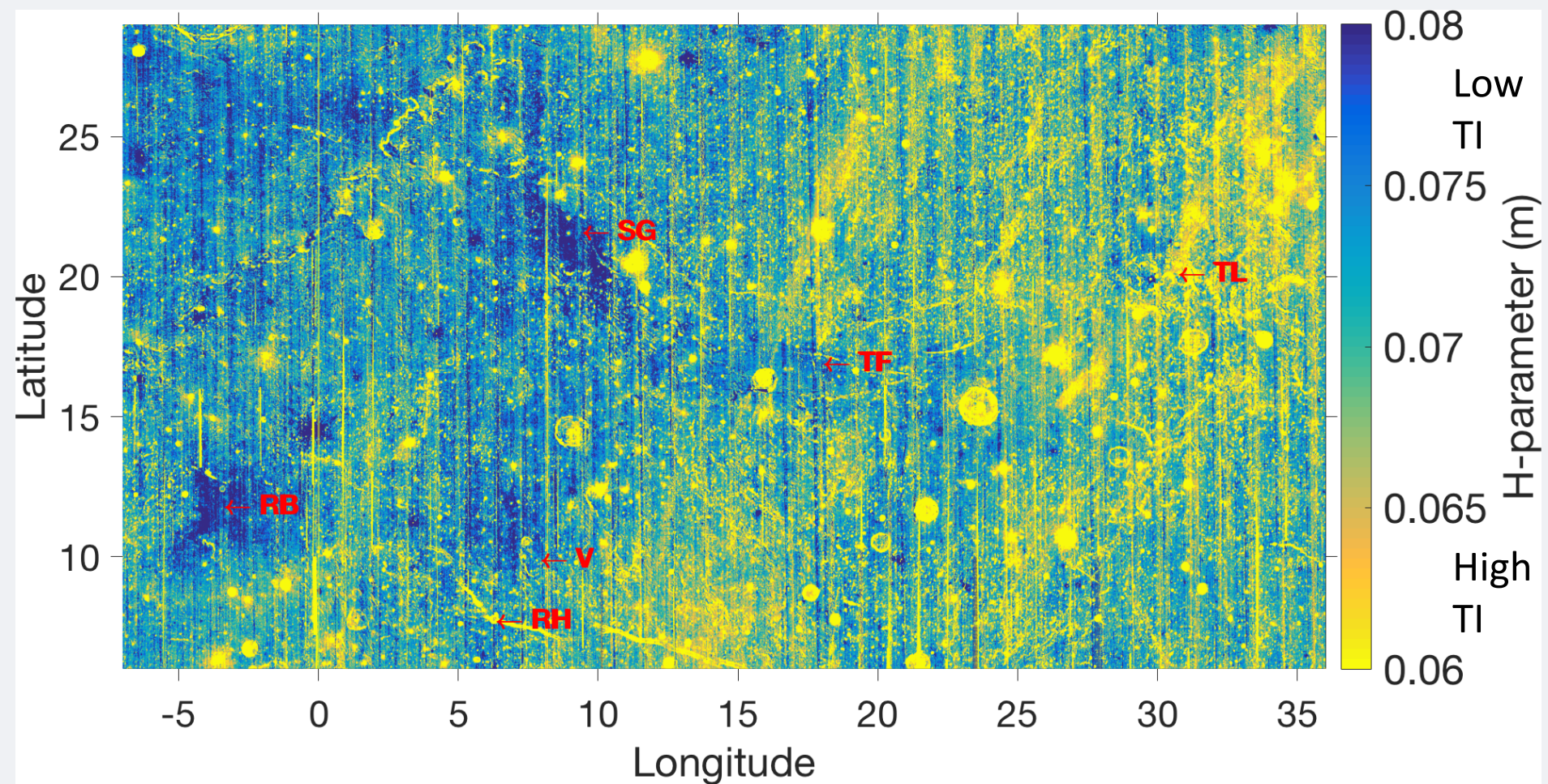
- Regional deposits ($>2500 \text{ km}^2$):
 - formed through lava-fountain eruptions
 - composed of crystalline beads and/or glass beads
- Localized deposits ($<2500 \text{ km}^2$):
 - formed through Vulcanian-style eruptions
 - incorporate a higher fraction of country rock
- Deposits of glass beads \rightarrow lower thermal inertia than regolith
- Country rock would increase the thermal inertia

Methods

- Fit Diviner nighttime regolith temperature measurements from Bandfield et al. (2011)
- Temperatures exclude rocks >1 m
- Thermal inertia proportional to density
- Assume regolith density:
$$\rho(z) = \rho_d - (\rho_d - \rho_s)e^{-z/H}$$
- Low $H \rightarrow$ high thermal inertia \rightarrow more rocks / lower porosity

(Hayne et al. submitted)



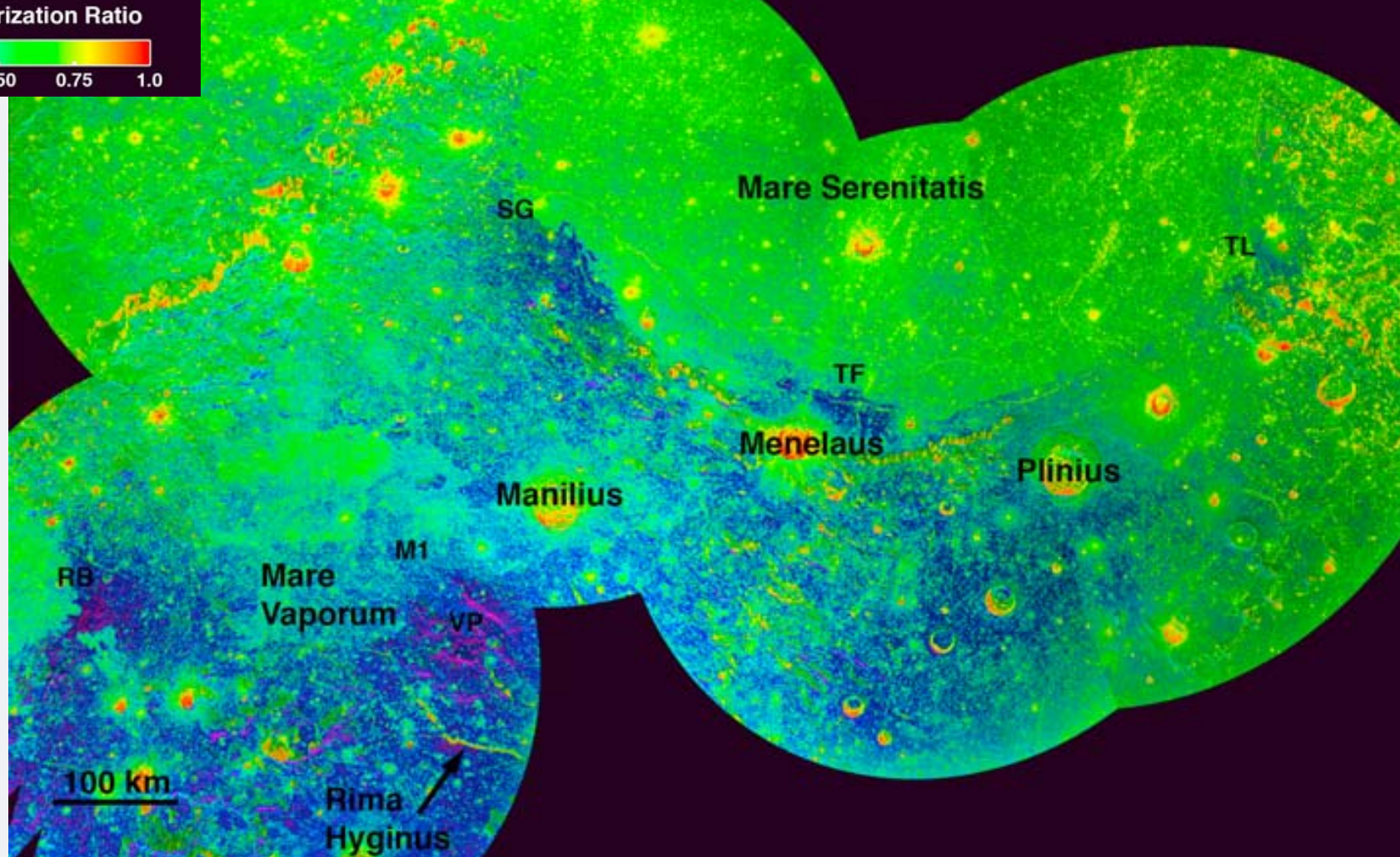


Rima Bode (RB), Sulpicius Gallus (SG), Vaporum (V), Taurus-Littrow (TL), Tacquet Formation (TF), Rima Hyginus (RH)

Circular Polarization Ratio



Carter et
al. (2009):
ground-
based
S-band
(12.6 cm)

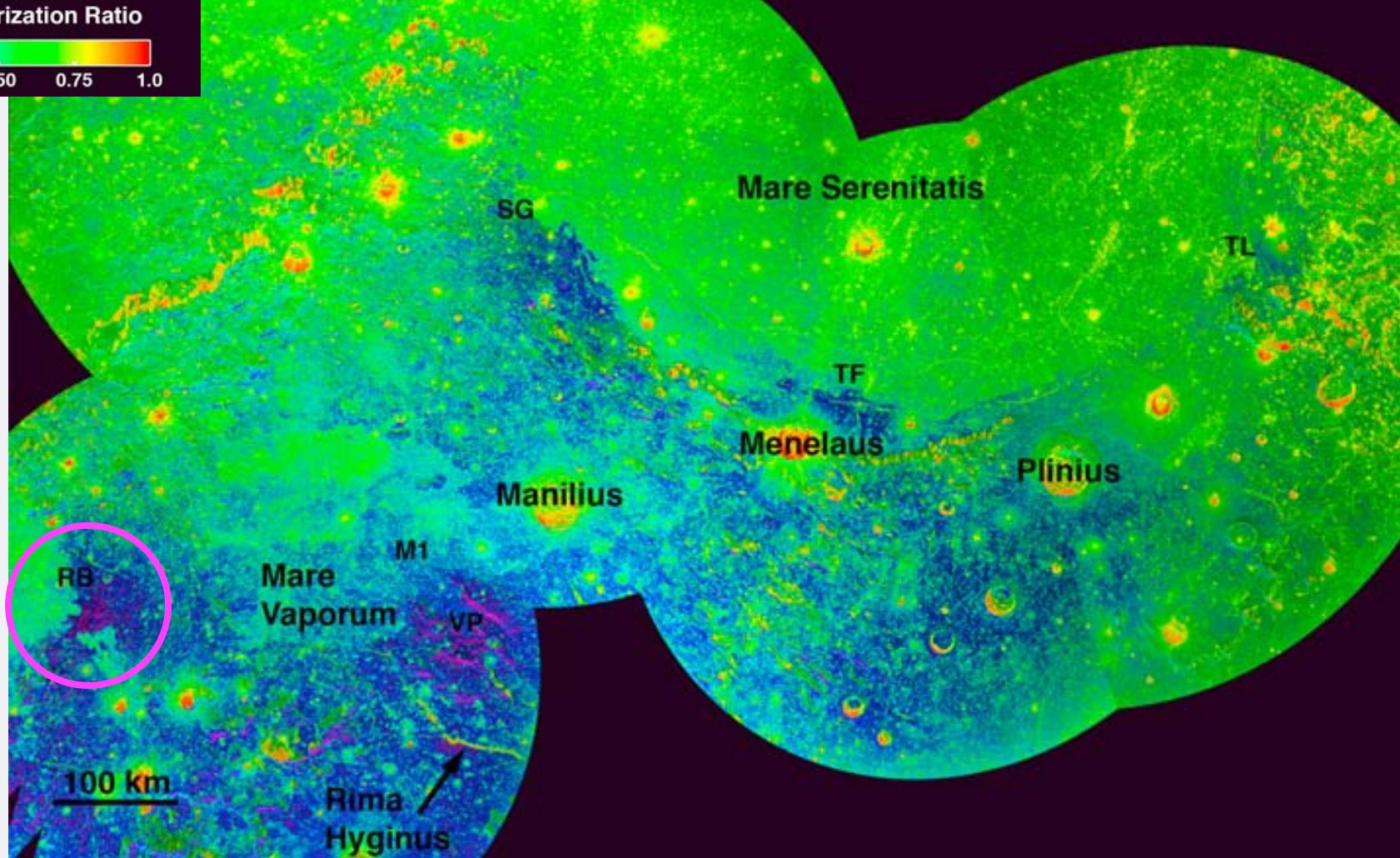


Rima Bode (RB), Sulpicius Gallus (SG), Vaporum (VP), Taurus-Littrow (TL), Tacquet Formation (TF)

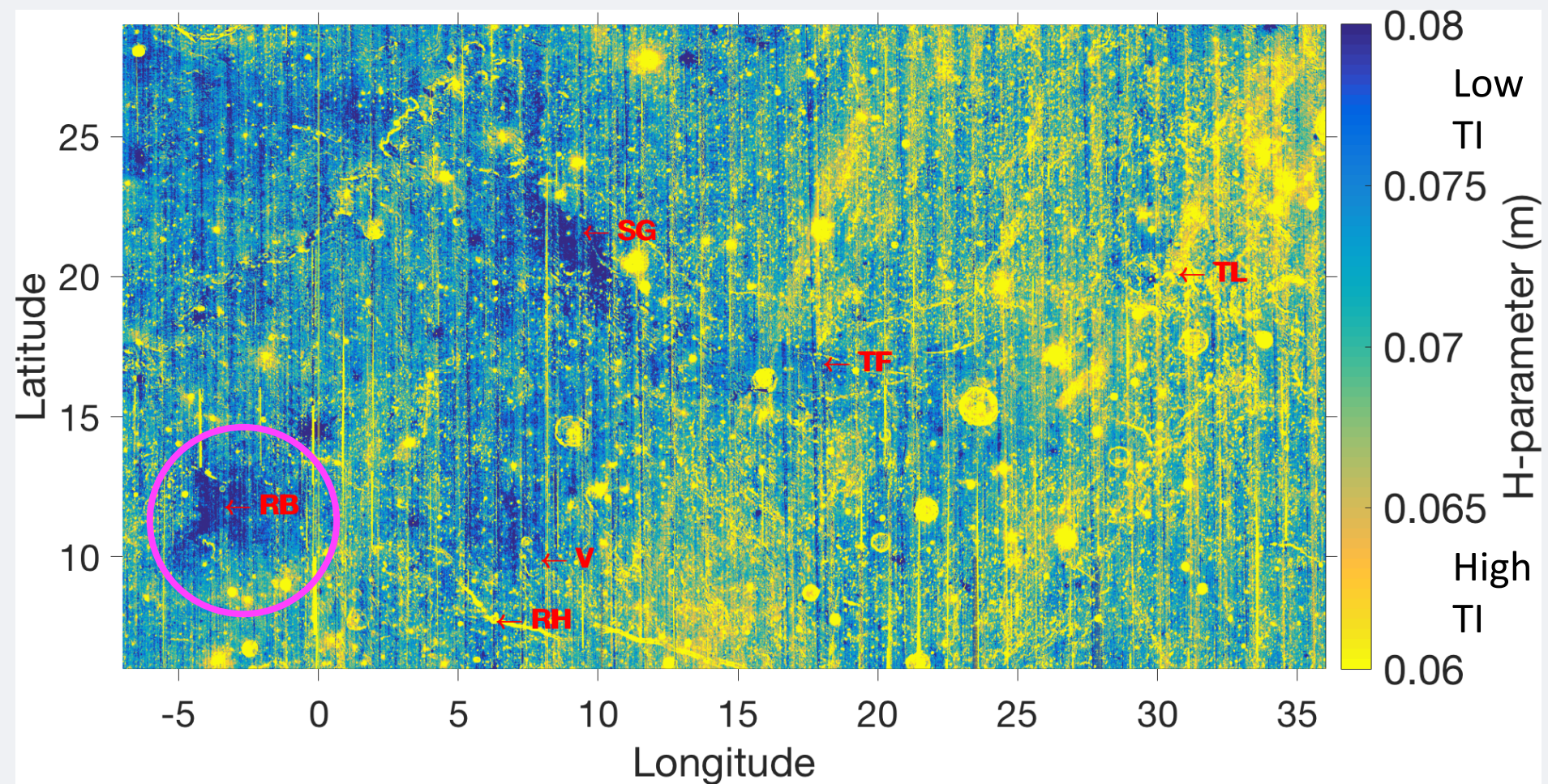
Circular Polarization Ratio



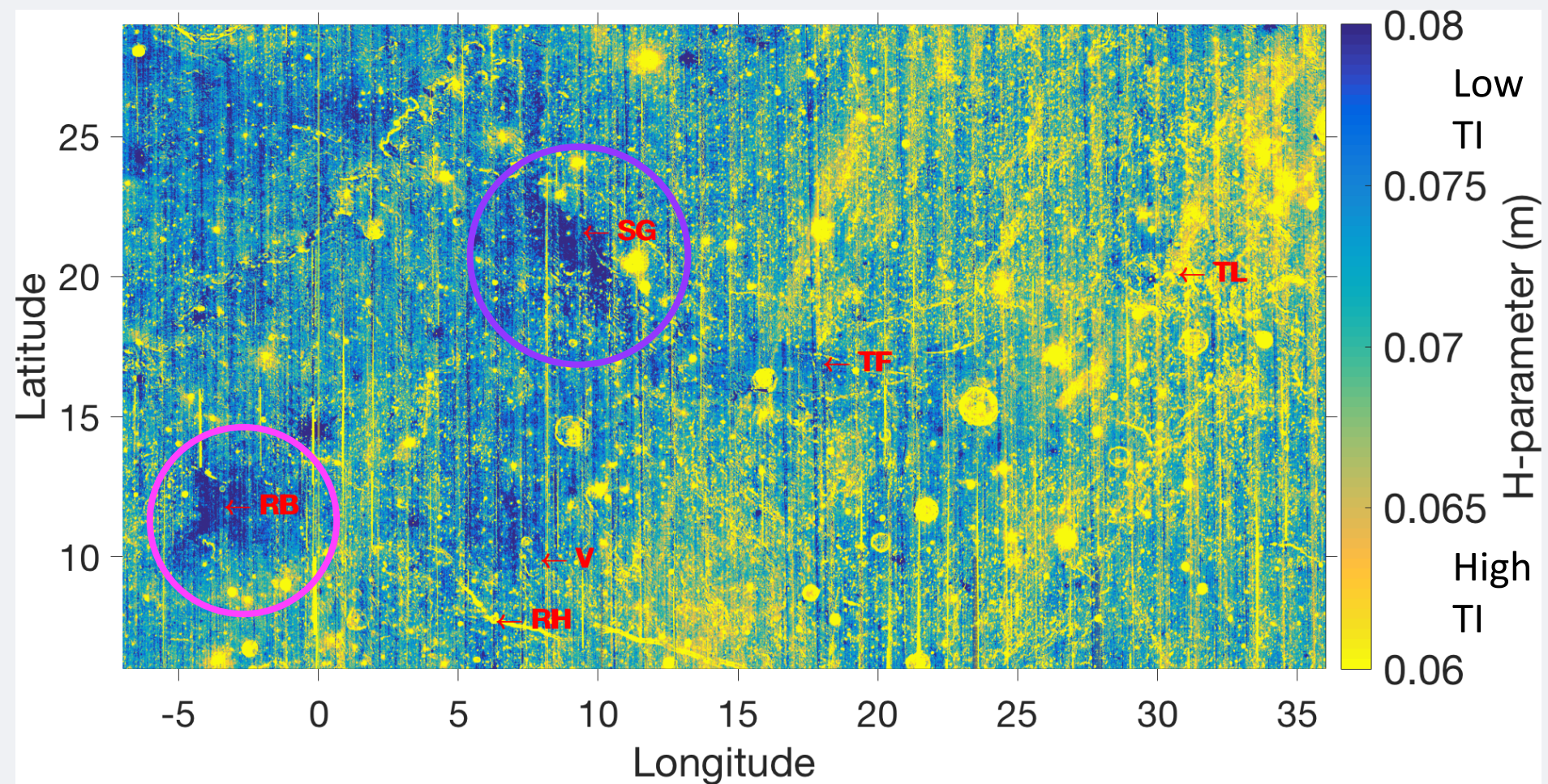
Carter et al. (2009):
ground-based
S-band
(12.6 cm)



Rima Bode (RB), Sulpicius Gallus (SG), Vaporum (VP), Taurus-Littrow (TL), Tacquet Formation (TF)



Rima Bode (RB), Sulpicius Gallus (SG), Vaporum (V), Taurus-Littrow (TL), Tacquet Formation (TF), Rima Hyginus (RH)

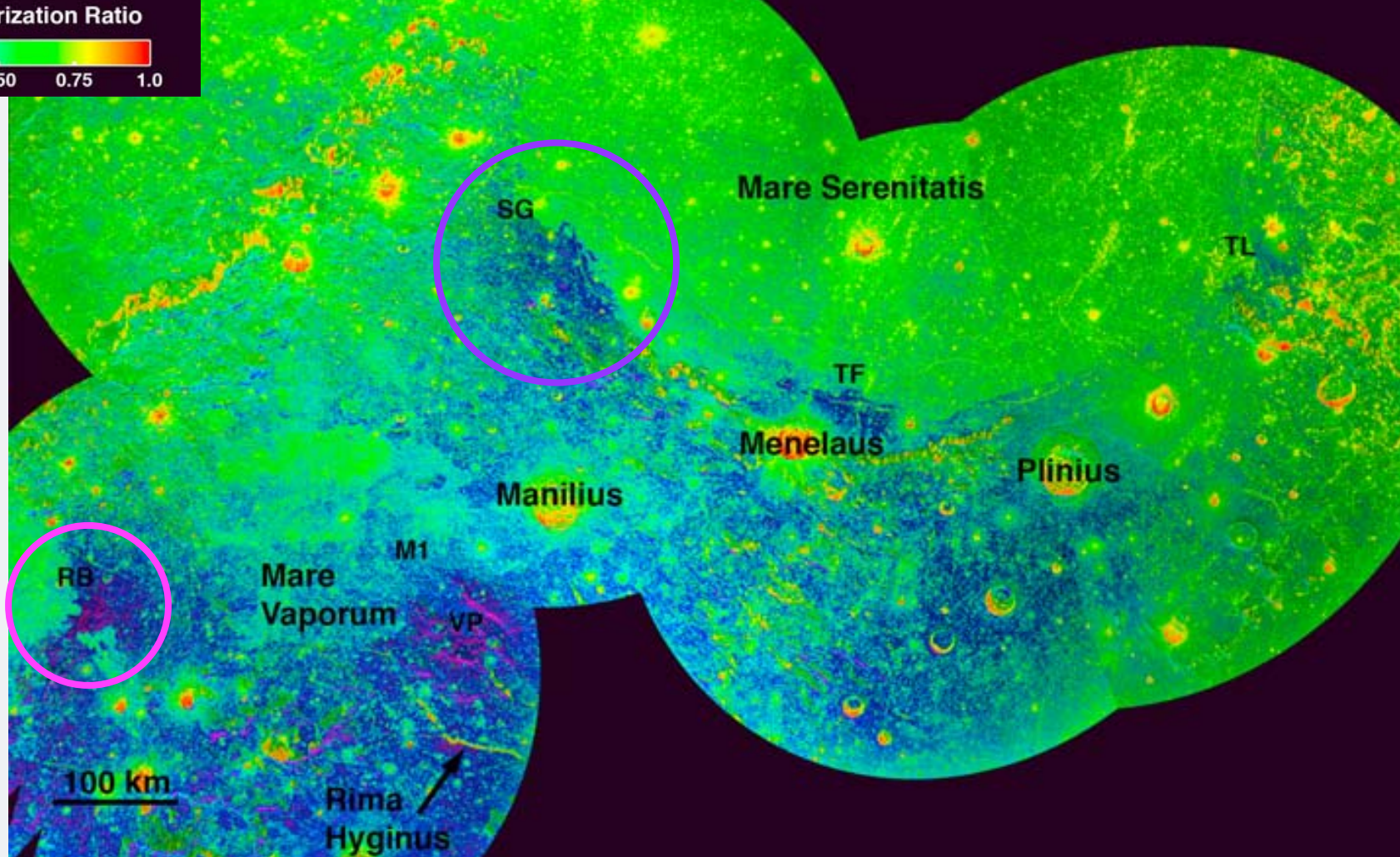


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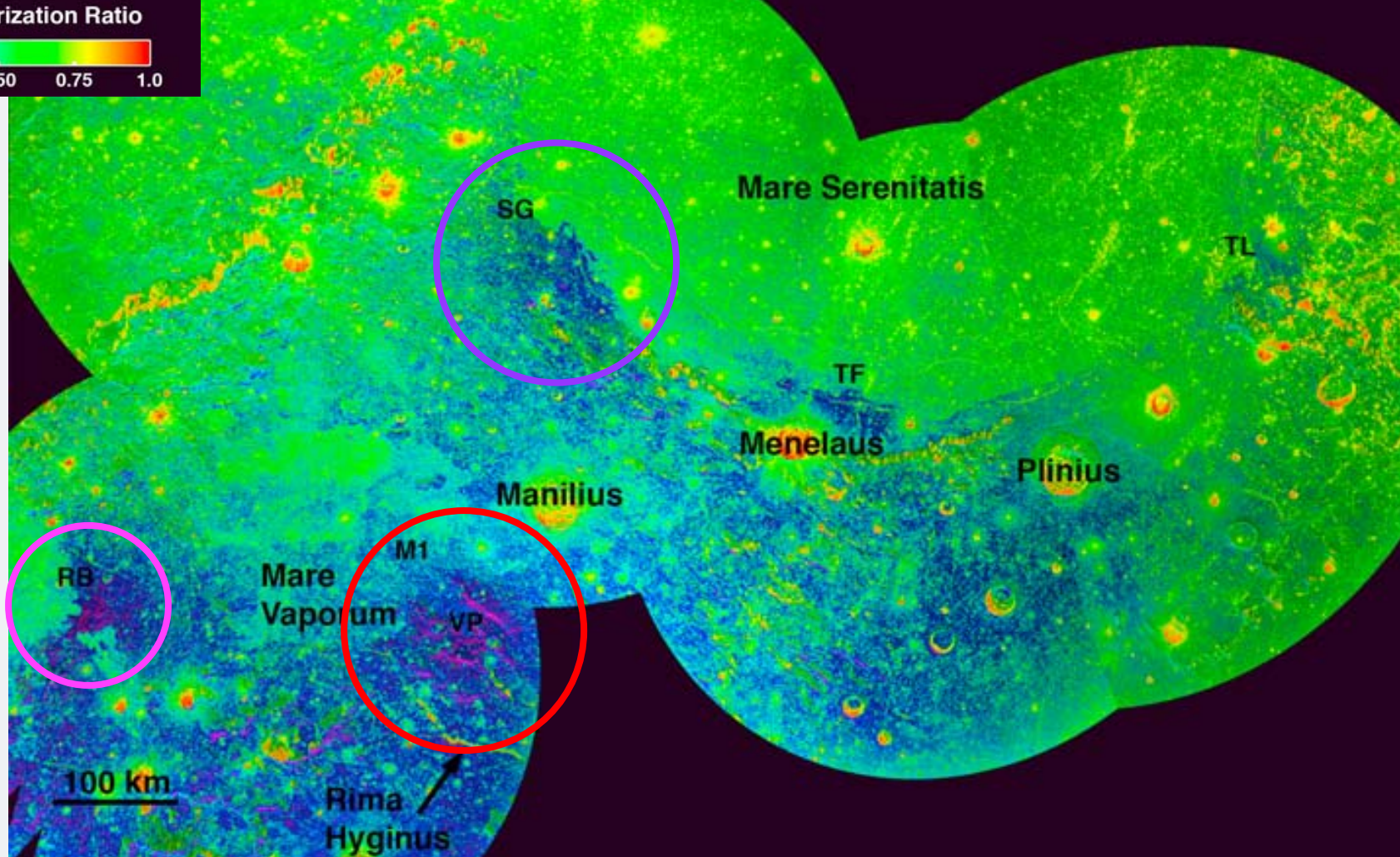


Rima Bode (RB), Sulpicius Gallus (SG), Vaporum (VP), Taurus-Littrow (TL), Tacquet Formation (TF)

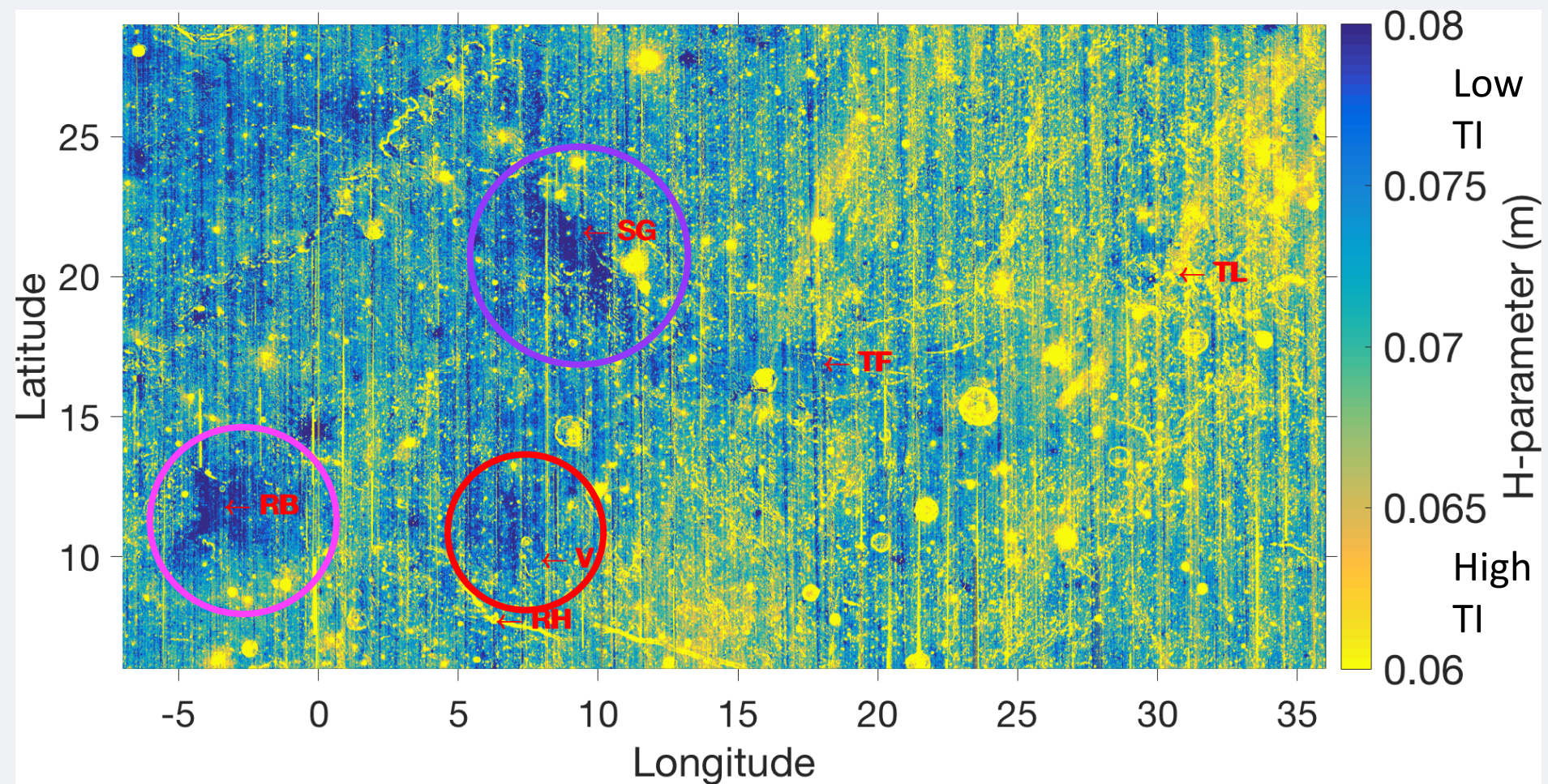
Circular Polarization Ratio



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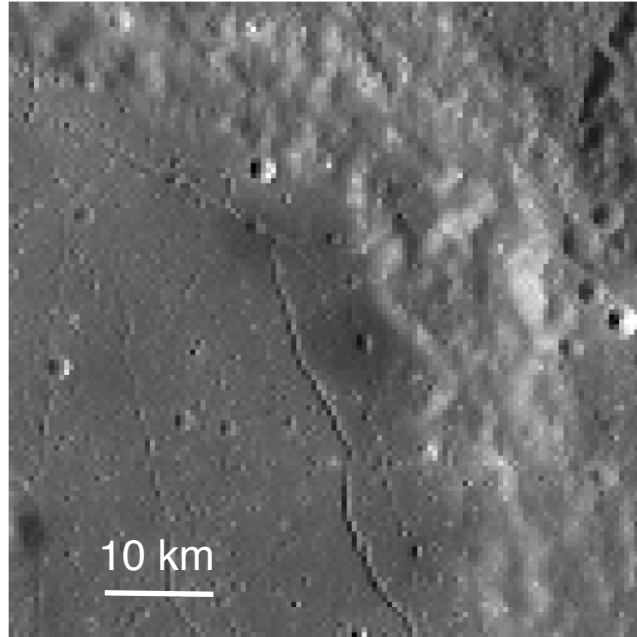
Rima Bode (RB), Sulpicius Gallus (SG), Vaporum (VP), Taurus-Littrow (TL), Tacquet Formation (TF)



Rima Bode (RB), Sulpicius Gallus (SG), Vaporum (V), Taurus-Littrow (TL), Tacquet Formation (TF), Rima Hyginus (RH)

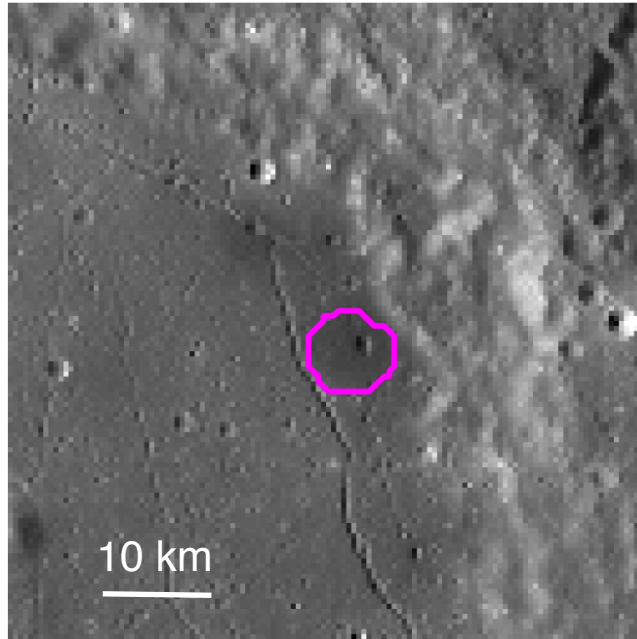
Localized Pyroclastic Deposits

Alphonsus NE

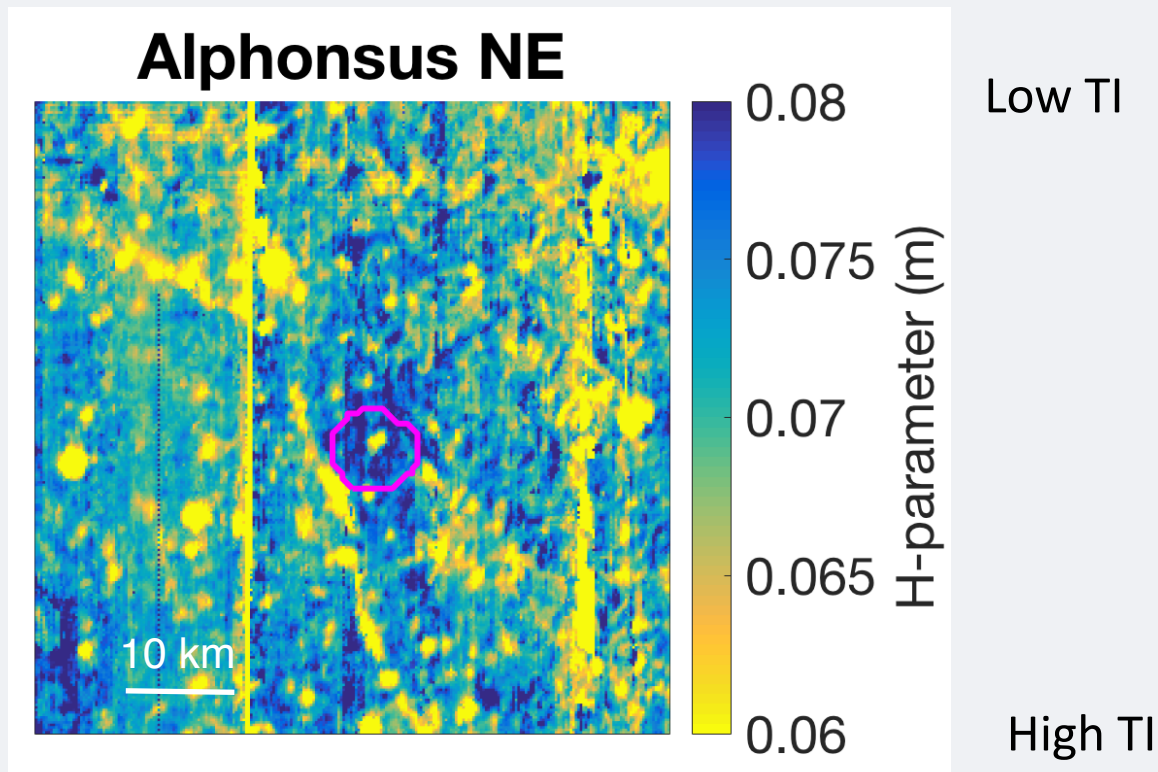


Localized Pyroclastic Deposits

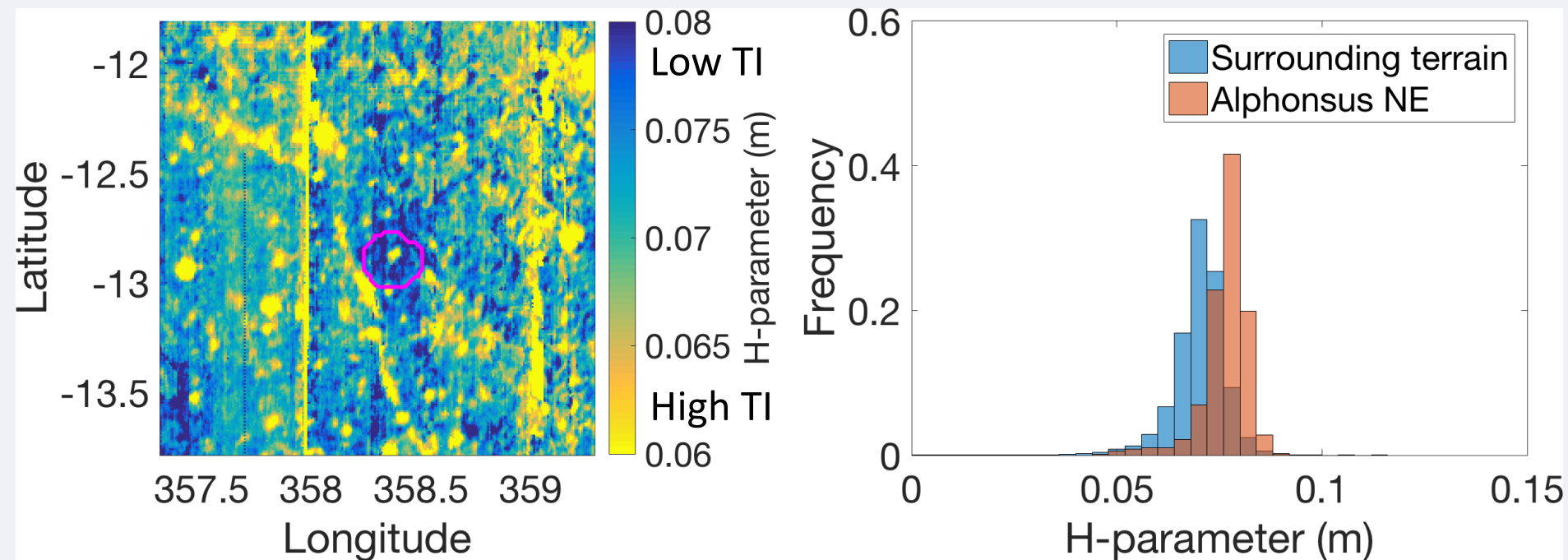
Alphonsus NE



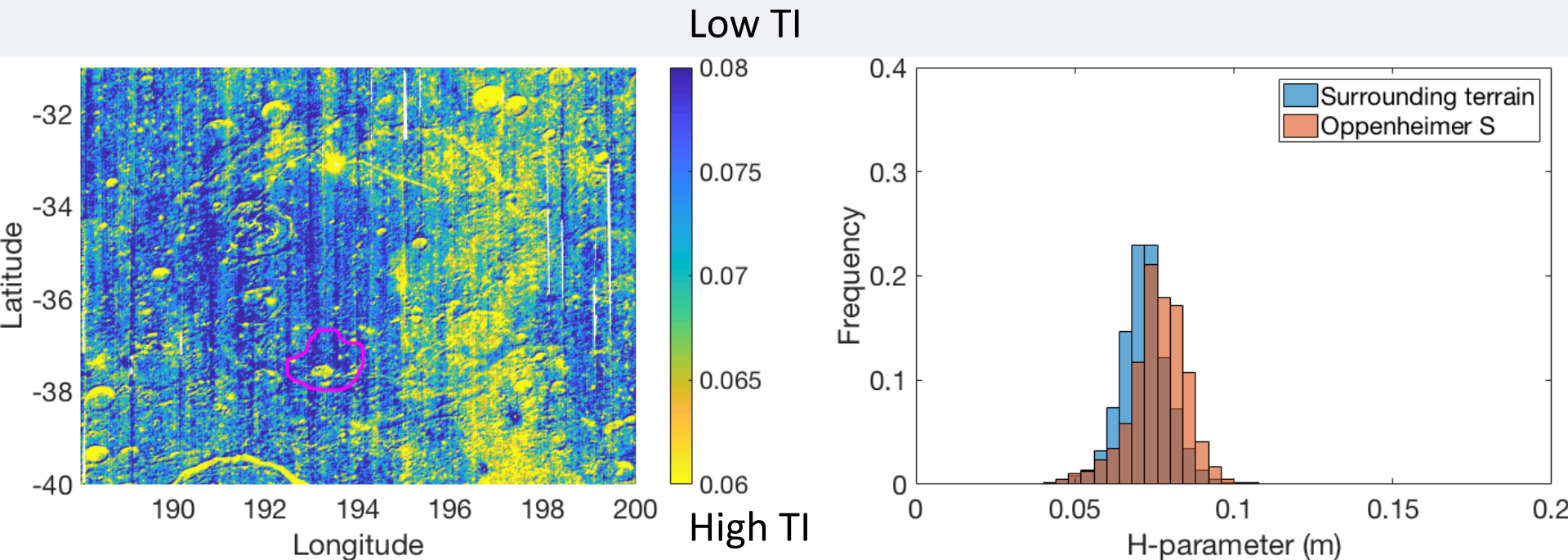
Localized Pyroclastic Deposits



Localized Pyroclastic Deposits



Localized Pyroclastic Deposits



Localized Pyroclastic Deposits

Lack of thermal inertia anomaly means:

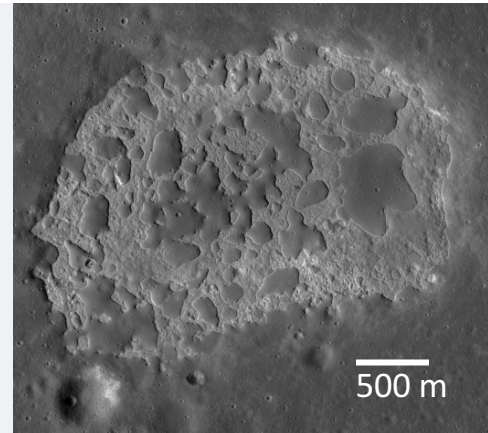
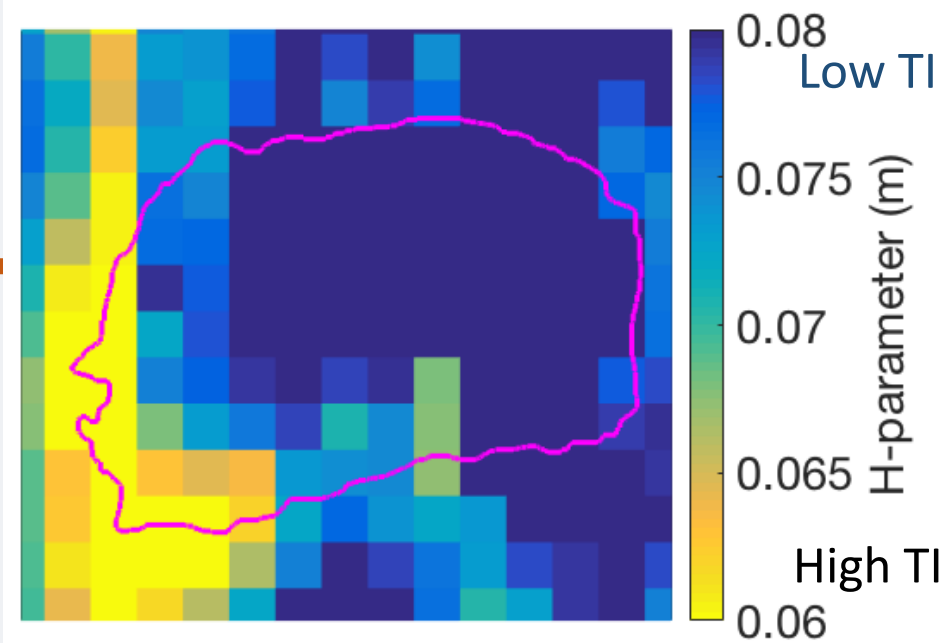
- a) Thin; easily punctured by impacts
- b) Small (in area); easily covered in rocks by nearby impacts or mass wasting of nearby slopes

— or —

- c) Original eruption incorporated a significant amount of country rock

Implications for Ina

- Magmatic foam
- Pyroclastic:
 - a) Ina is younger than most localized deposits
 - b) Deposit at Ina is thicker
 - c) Deposit at Ina incorporated less country rock



Conclusions

- Regional pyroclastic deposits have a lower thermal inertia than regolith
- Some localized pyroclastic deposits also have a lower thermal inertia than regolith, but some do not
 - Difficult to separate effects of eruption conditions from post emplacement modification
- Ina has a low thermal inertia consistent with either magmatic foam or young pyroclastic deposits

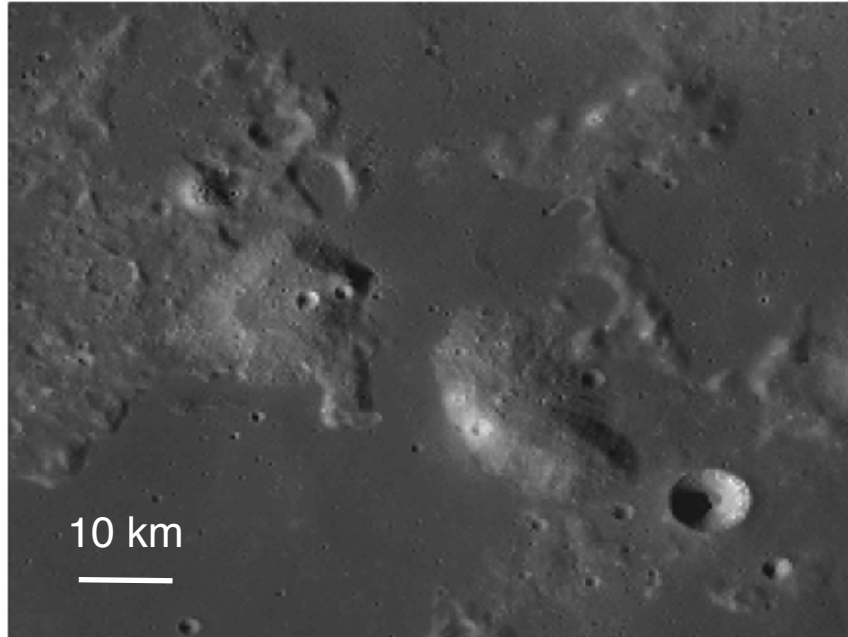
Back up slides

Lunar 'Red Spots'

- Spectrally red
- Domes (with steep slopes and high albedos), smooth plains, shields, or rugged patches of highland material
- Generally pre-date mare volcanism; formed over shorter interval
- Created by viscous lava similar to terrestrial dacites, basaltic andesites, or rhyolites
- Some are highly silicic

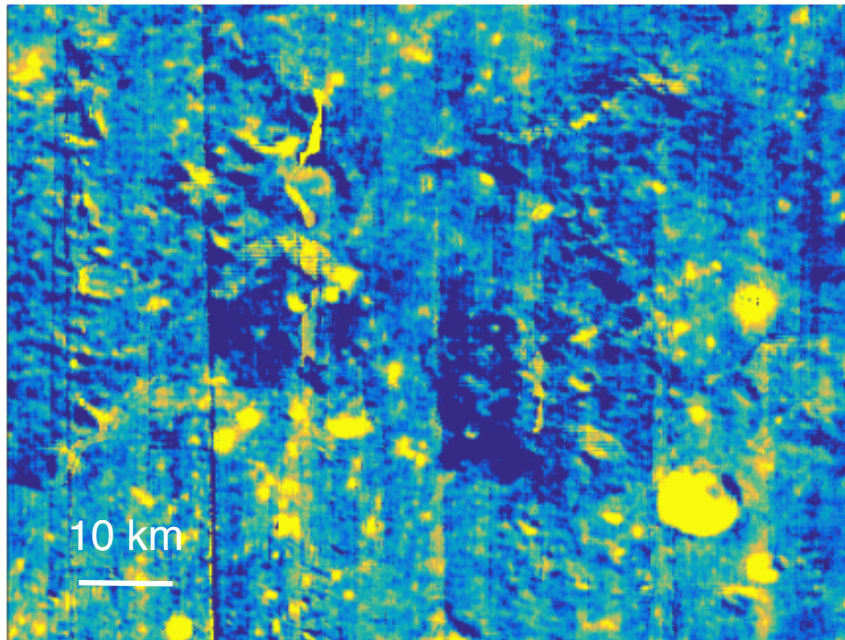
Lunar Red Spots

Gruithuisen Domes



Lunar Red Spots

Gruithuisen Domes



10 km

0.08

0.075

0.07

0.065

0.06

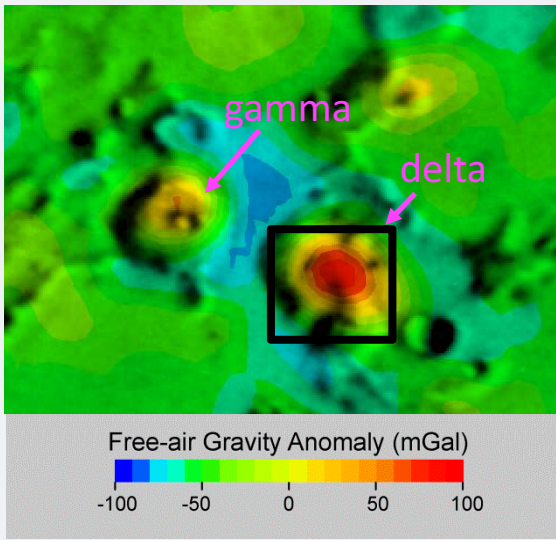
H-parameter (m)

Low TI

High TI

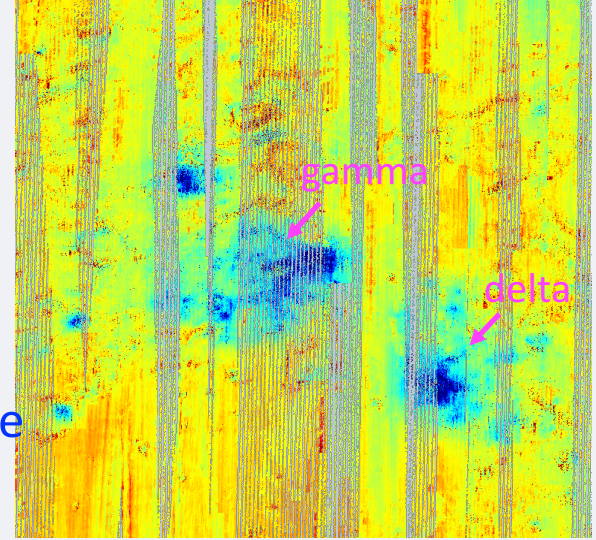
Free-air
Gravity
Anomaly
(Kiefer et
al., 2016):

Bulk density
= 2150 kg m^{-3}

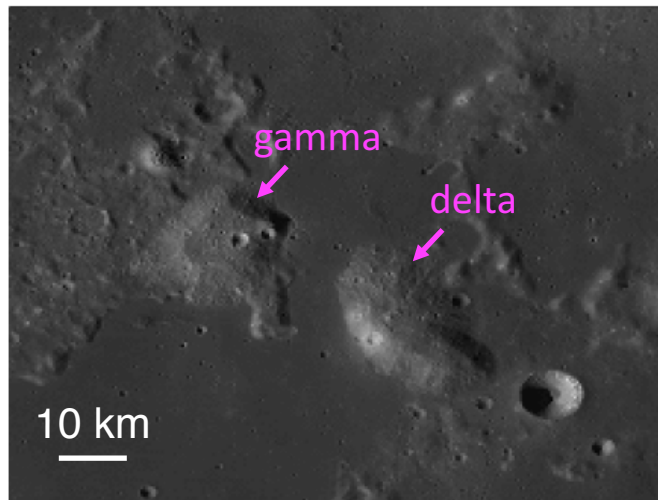


Christiansen
Feature (CF)
map:

Blue = more
silicic



LROC
WAC:



Thermal
Inertia:

Blue = low
TI

